



*ASTROPHYSICAL
TERAHERTZ
LARGE
AREA
SURVEY*

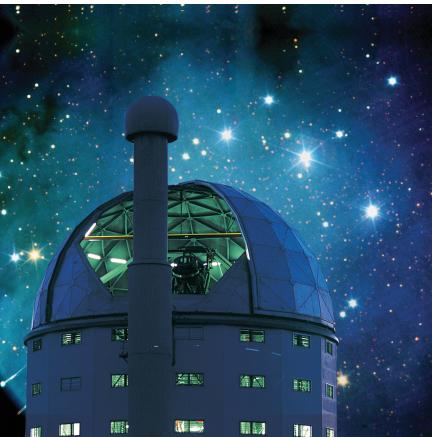


SALT Spectra of Gravitationally Lensed Herschel Systems

Lerothodi L. Leeuw

University of South Africa
in collaboration with the H-ATLAS team
and SALT Partner Members!!

(<http://www.h-atlas.org/>)



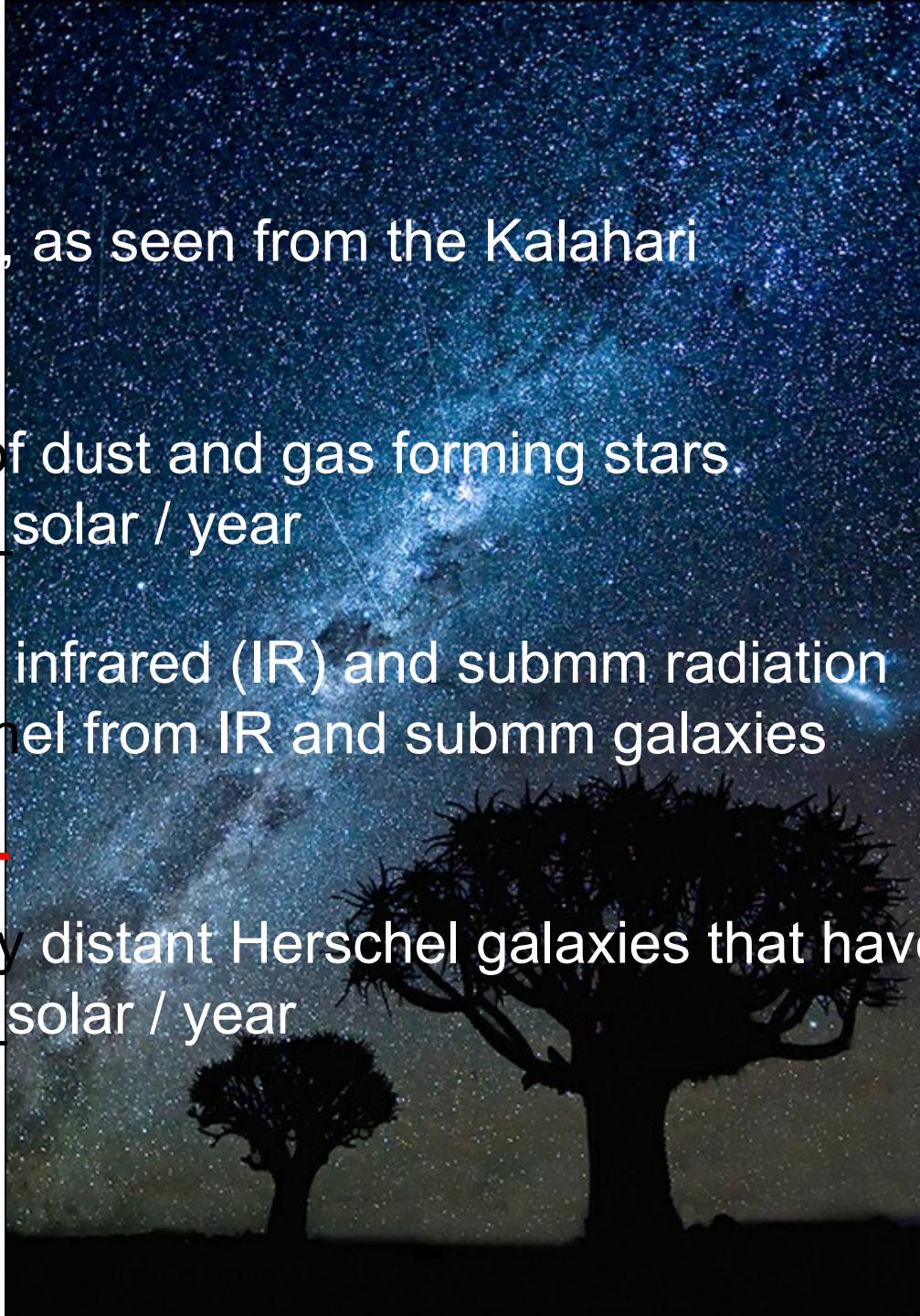
**Molagodimo,
the Milky Way**, as seen from the Kalahari
in the optical.

Dark patches of dust and gas forming stars
At 0.5 to 10 M_{solar} / year

The dust emits infrared (IR) and submm radiation
seen by Herschel from IR and submm galaxies

Optical / SALT

Lenses magnify distant Herschel galaxies that have
upto ~1000 M_{solar} / year



radio continuum (408 MHz)

atomic hydrogen

radio continuum (2.5 GHz)

molecular hydrogen

infrared

mid-infrared

near infrared

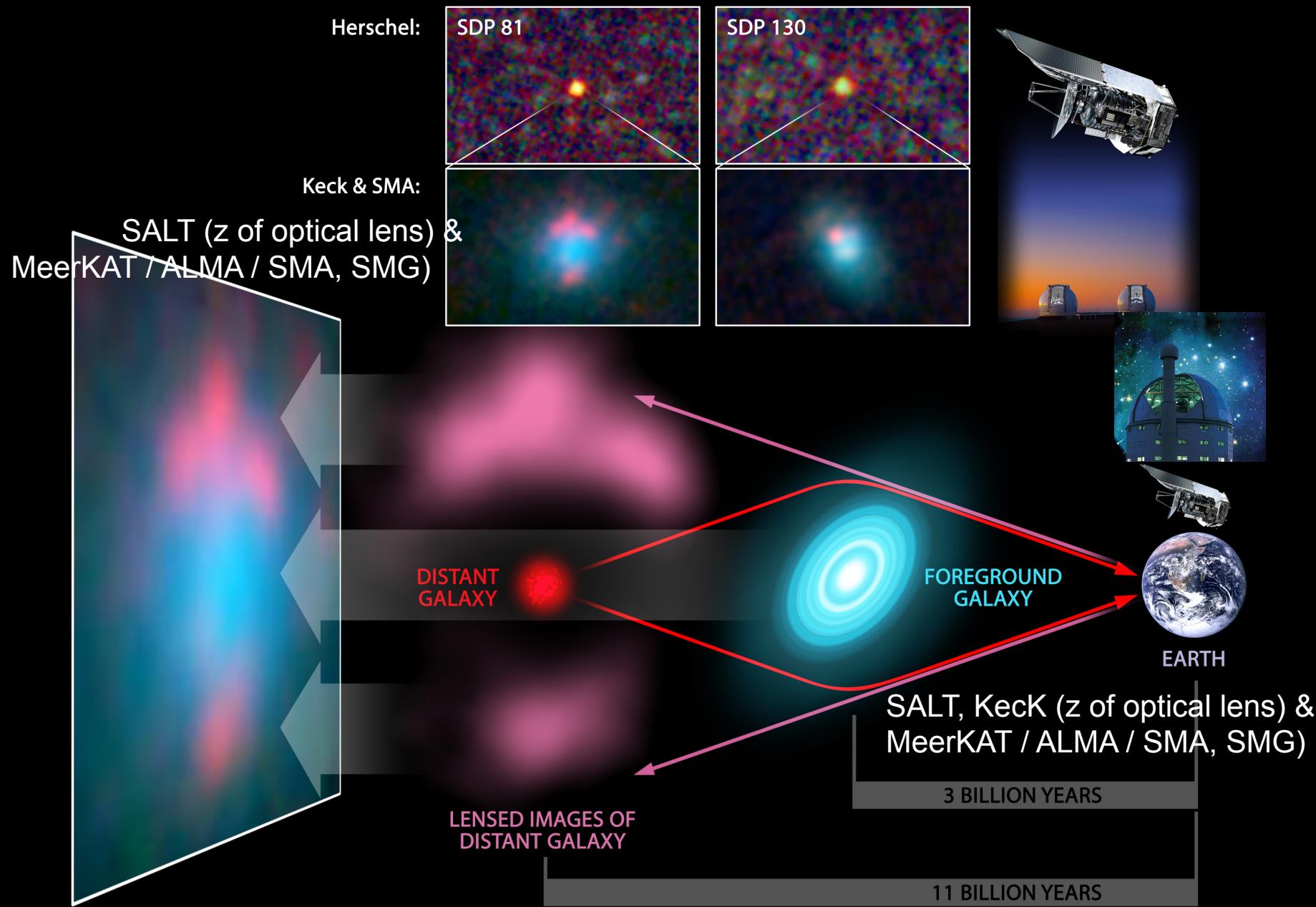
optical

x-ray

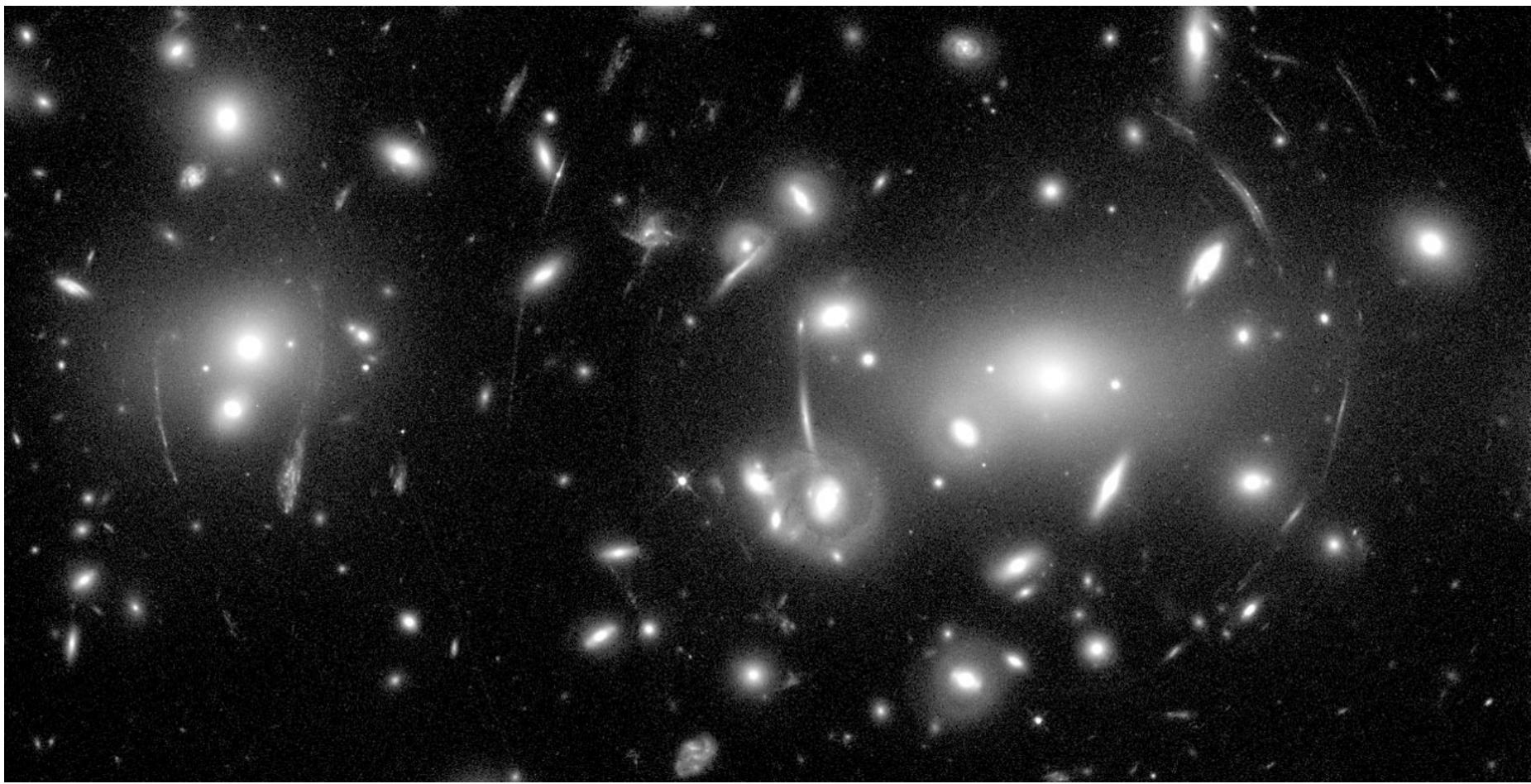
gamma ray



Multiwavelength Milky Way



CREDITS: ESA/NASA/JPL---Caltech/Keck/SMA



Hubble Space Telescope Image of Abell 2218



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The Hershel-ATLAS

The Hershel
Space
Observatory

- Largest Open Time Key Project on Herschel (Eales et al. 2010)
- 600 hours for about 550 deg^2 in N, S, and Equatorial fields
- PACS+SPIRE: 110, 160, 250, 350, and $500\mu\text{m}$
 - $\sim 9'', \sim 13'', \sim 20'', \sim 30'', \text{ and } \sim 40''$
- Led by Steve Eales (Cardiff) & Loretta Dunne (Nottingham + NZ) + 150 co--Is

A LENSING SCIENCE CASE FOR H-ATLAS

Sub-mm surveys are ideal for finding lenses

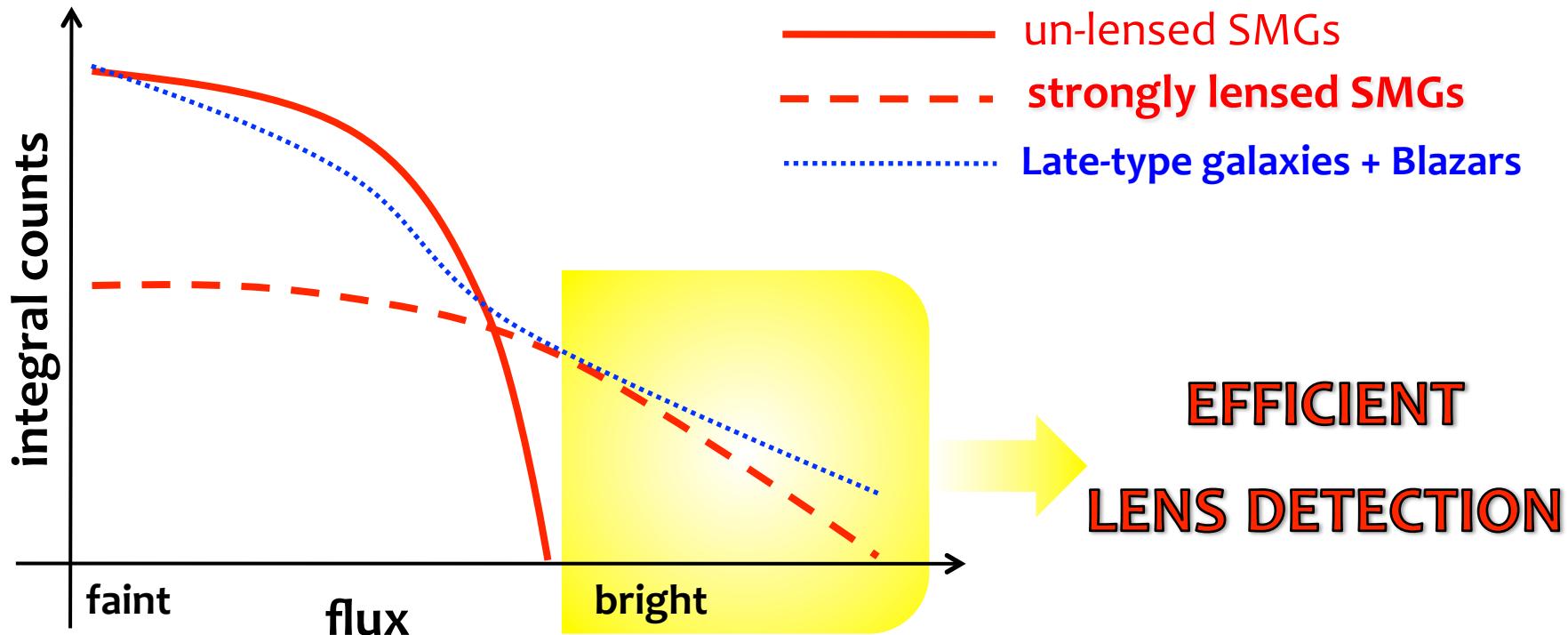
Blain (1996), Perrotta et al. (2003), Negrello et al. (2007, 2010)

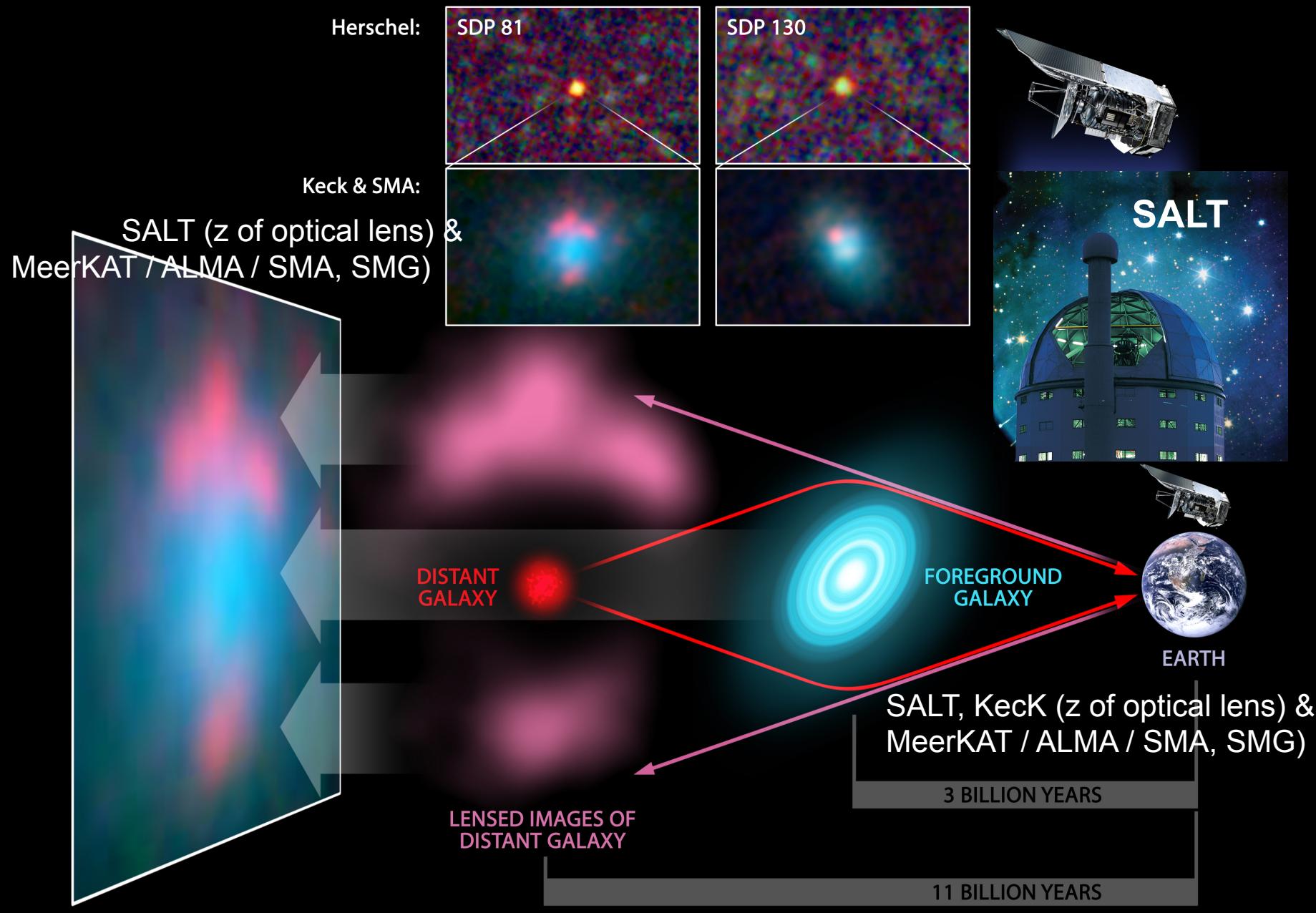
- **high redshift** → **high efficiency for lensing**

Chapman et al. (2005)

- **steep counts** → **strong magnification bias**

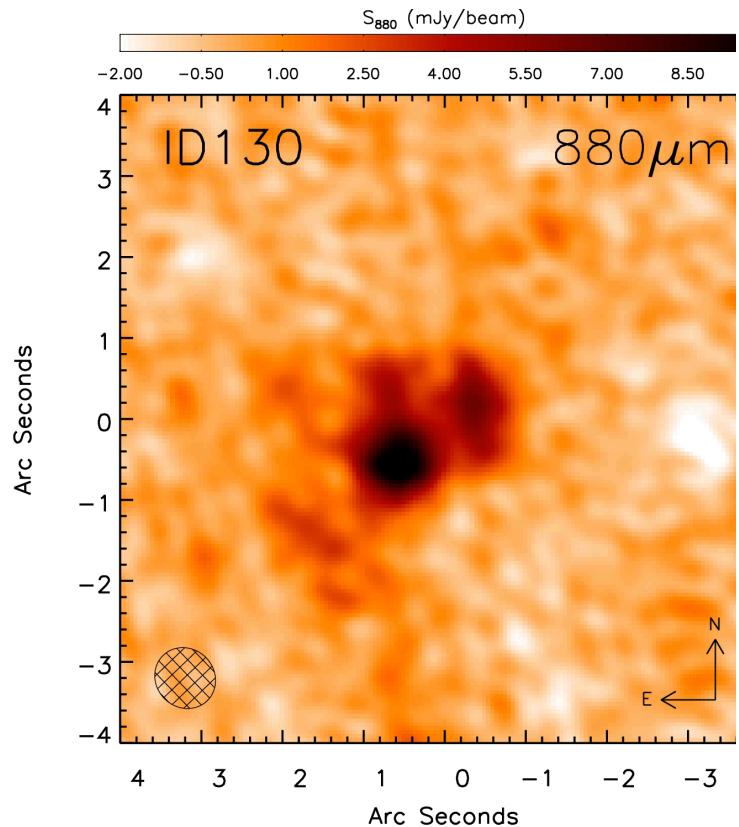
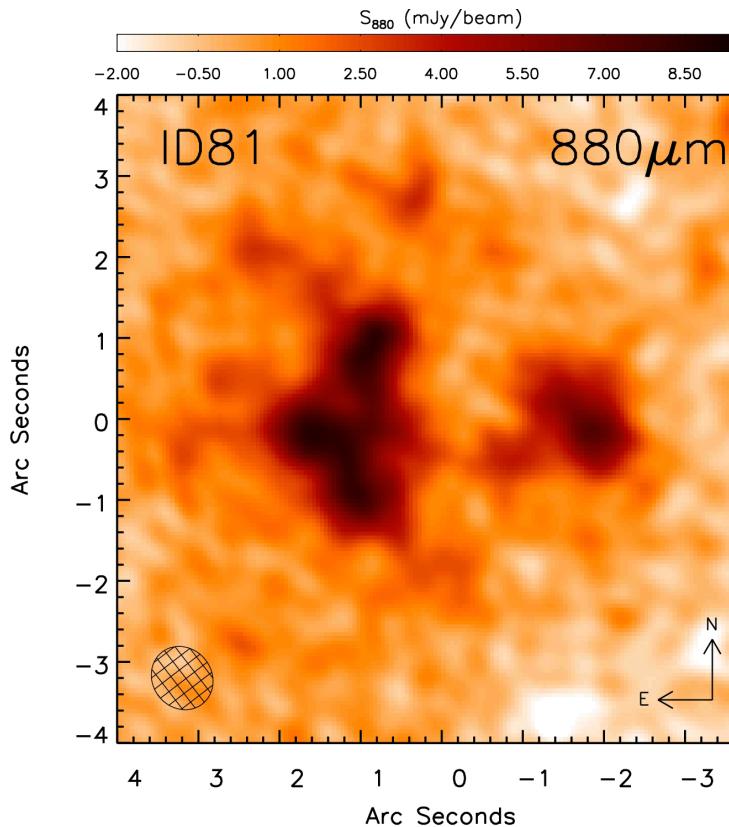
Coppin et al. (2006)





CREDITS: ESA/NASA/JPL---Caltech/Keck/SMA

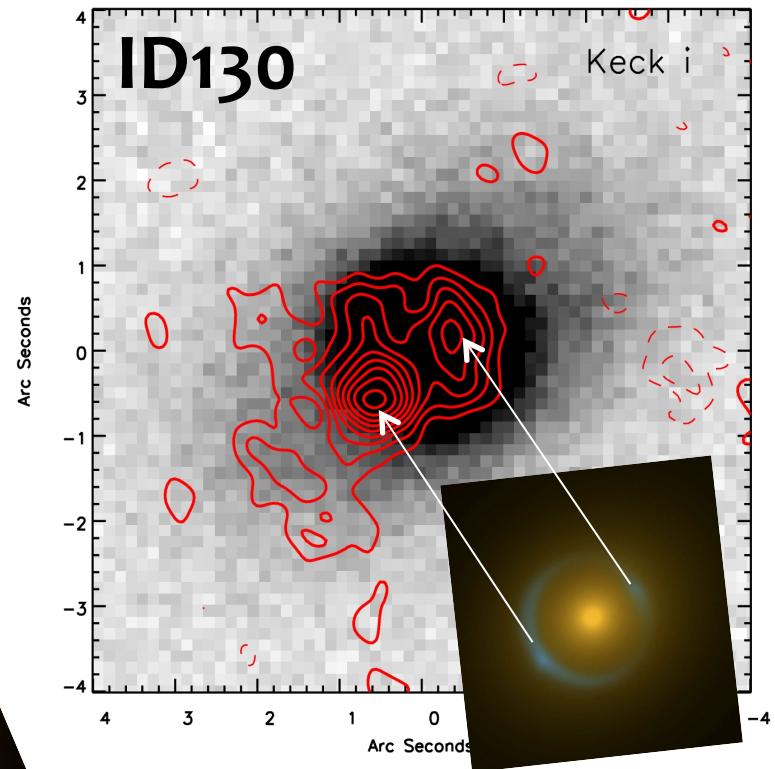
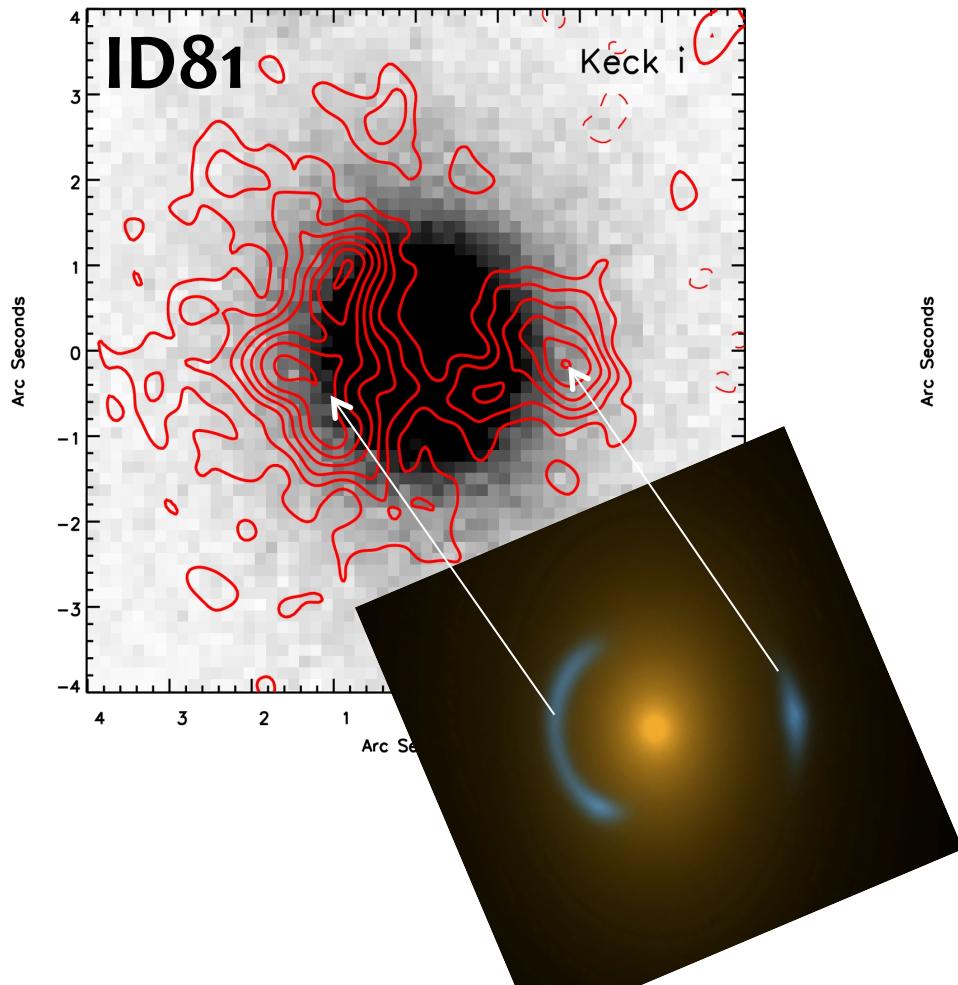
Evidence for Gravitational Lensing – Multiple Peaks in High Resolution Submm Imaging with the SMA, Negrello et al. 2010



GRAVITATIONAL LENS CANDIDATES ID81 – ID130

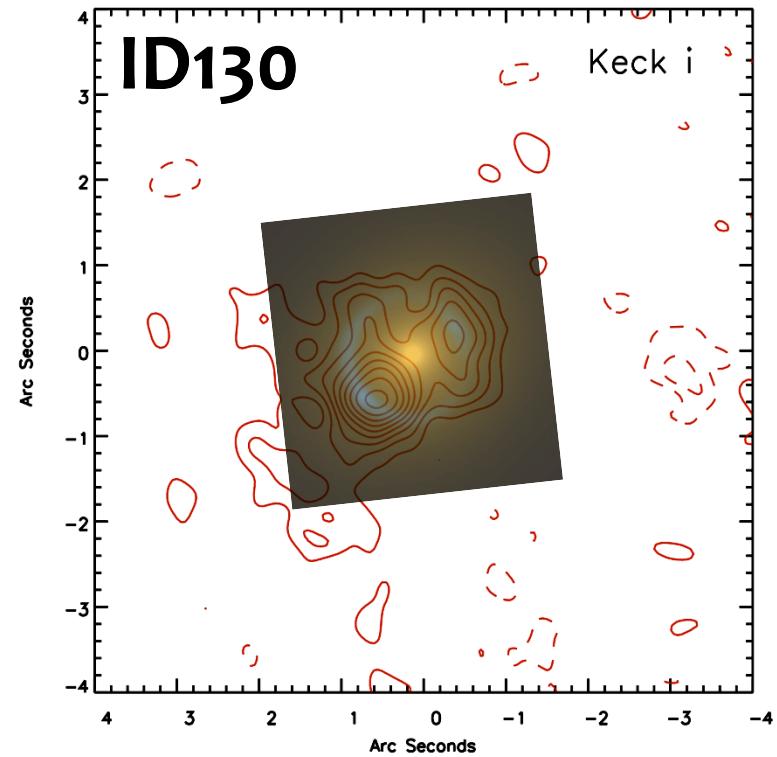
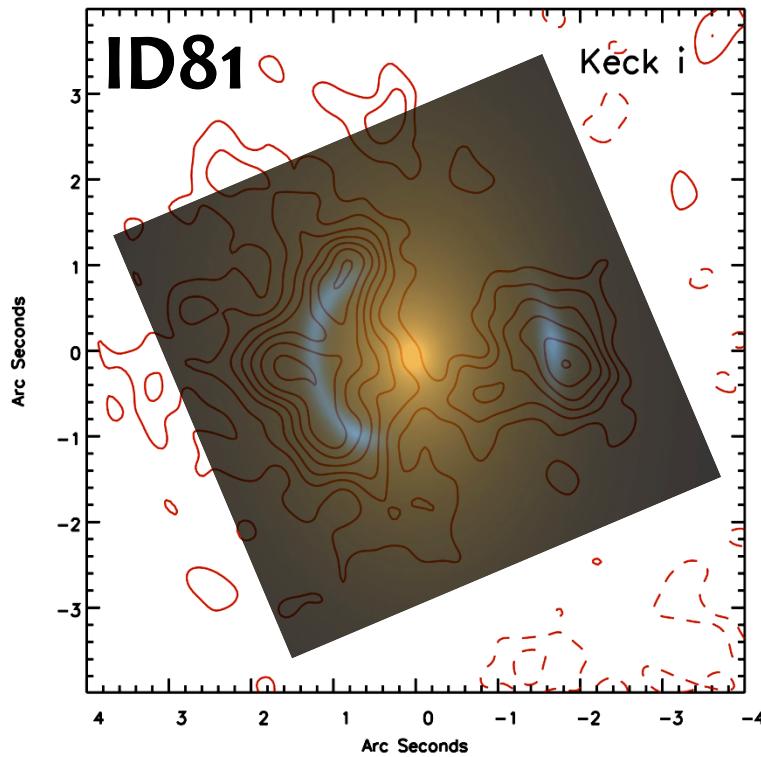
Sub Millimeter Array follow-up at 870 μ m

(very-extended, sub-compact and compact configurations)

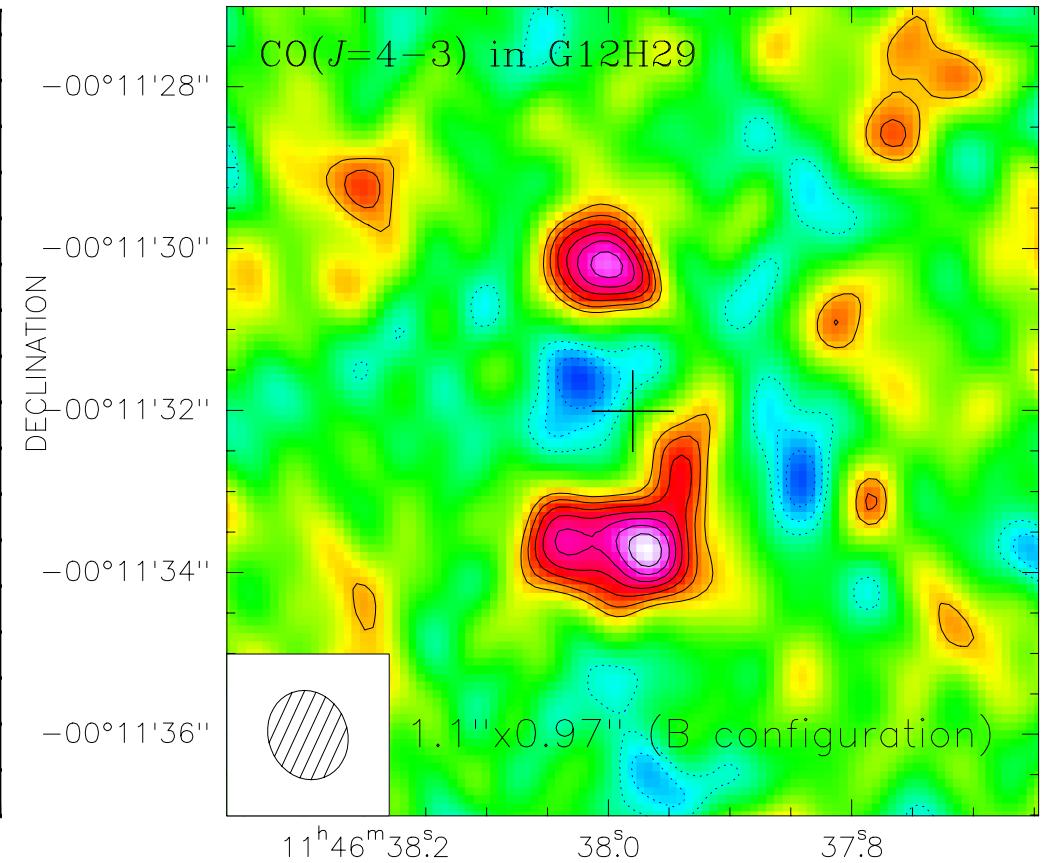
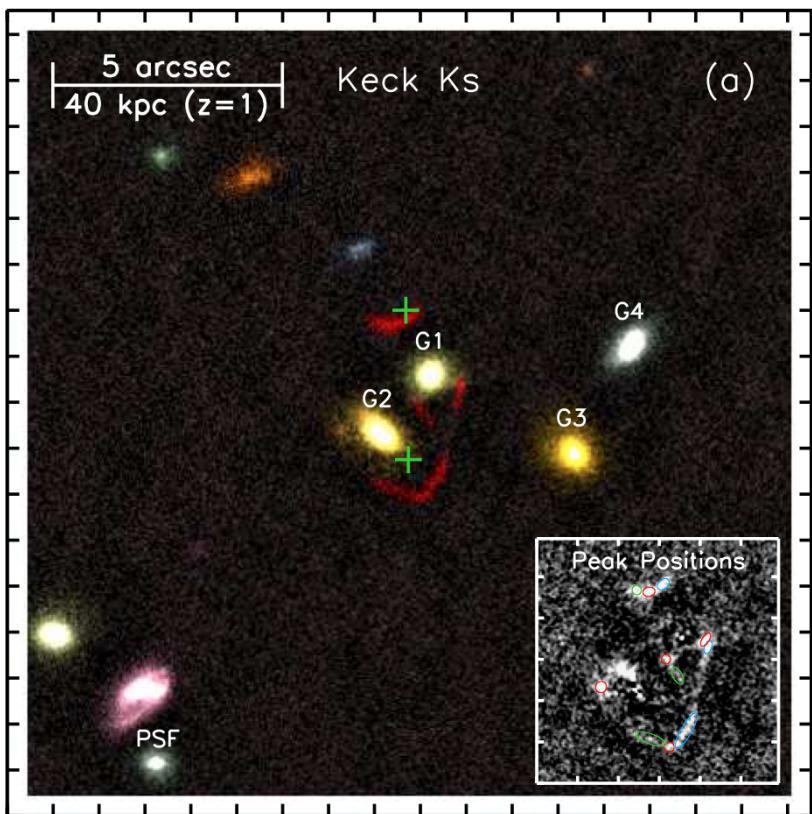


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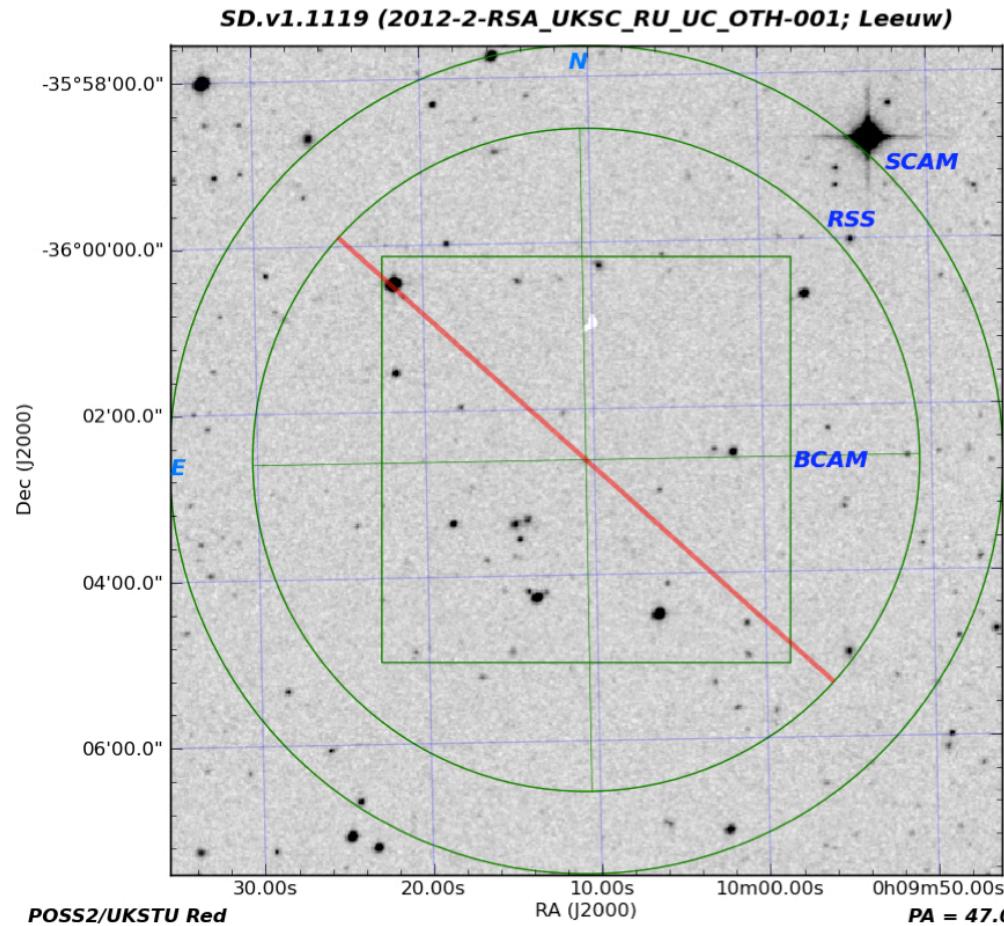
KECK Imaging and CARMA CO Maps of H-ATLAS G12H29 (z=3.259), a possible proto-Galaxy also detected with Planck



Left: Fu et al 2011. Right: Leeuw et al. in prep

SALT: Lens Redshift + MeerKAT: Masers + Conti + >> z CO

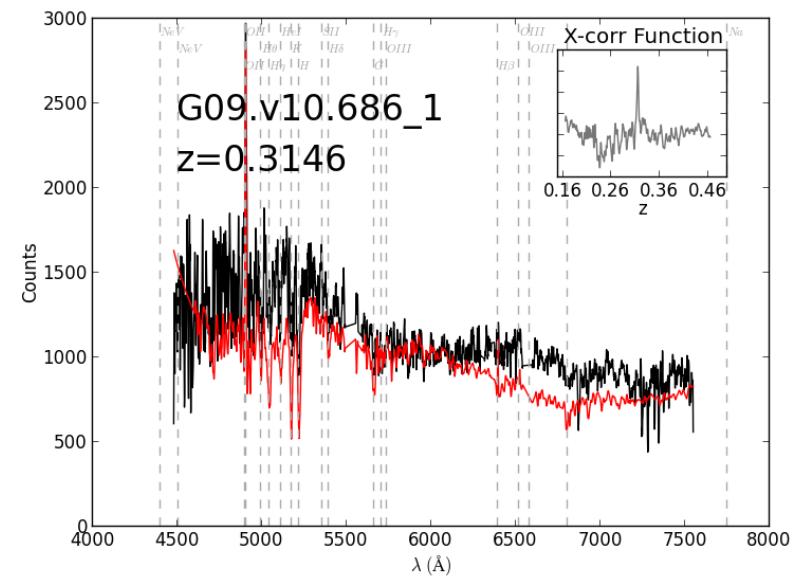
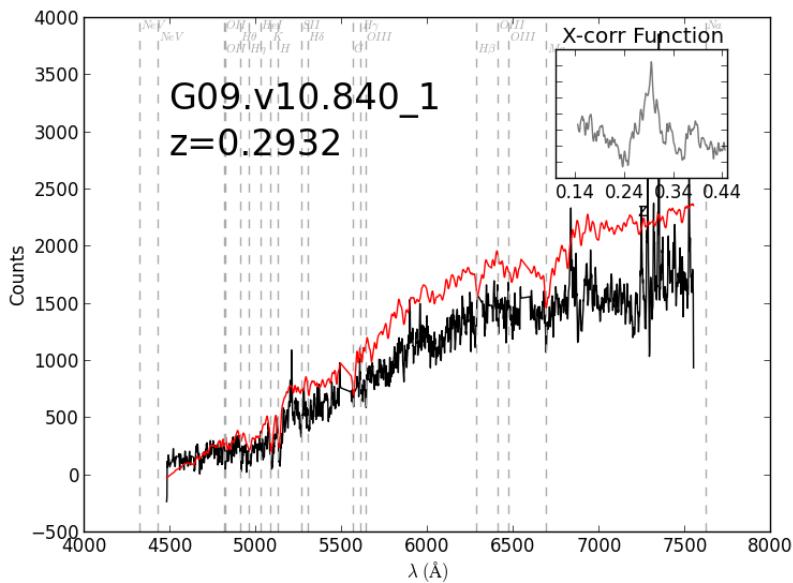
Typical SALT Finding Chart HATLAS Lens SDv1.1119 (V~20) with $z = 0.443$



Leeuw, Crawford, H-ATLAS-team et al. SAIP 2013



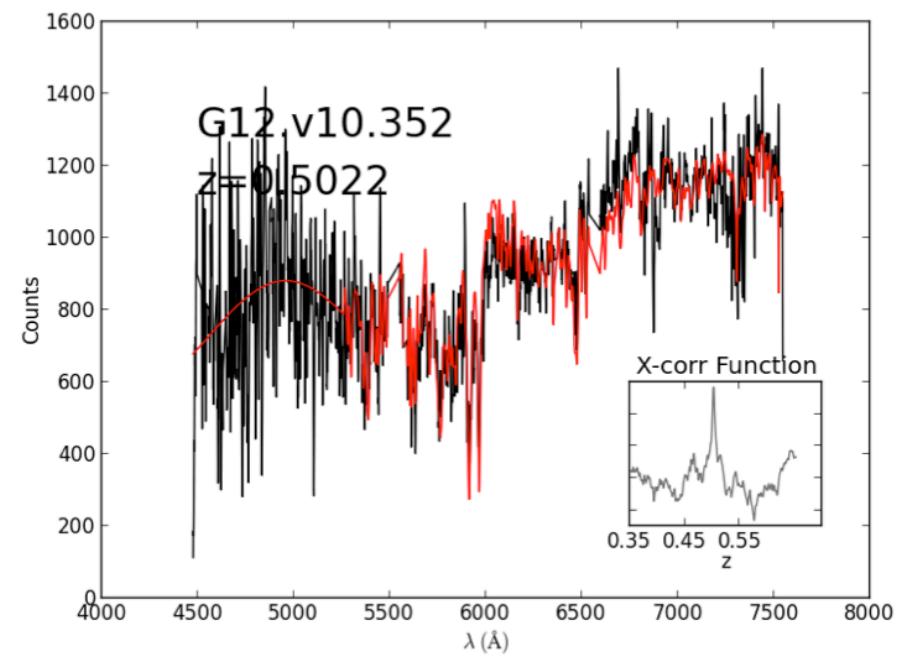
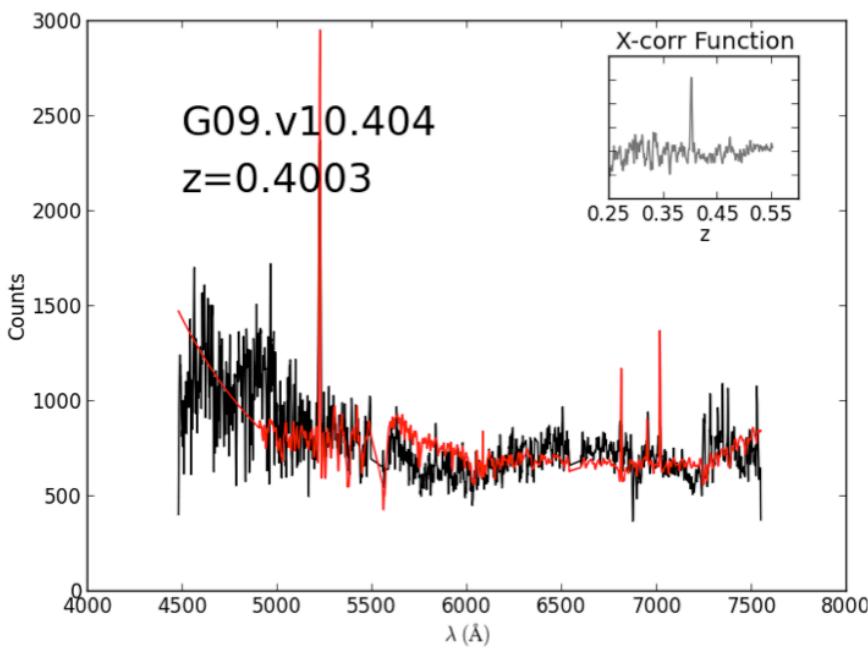
SALT Detected Spectra and Determined Redshifts



Leeuw, Crawford, H-ATLAS-team et al. SAIP 2013



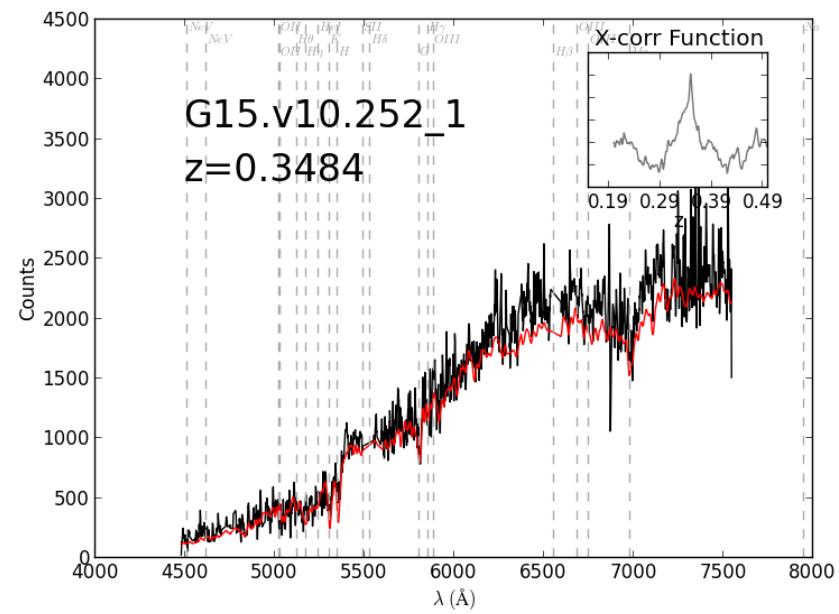
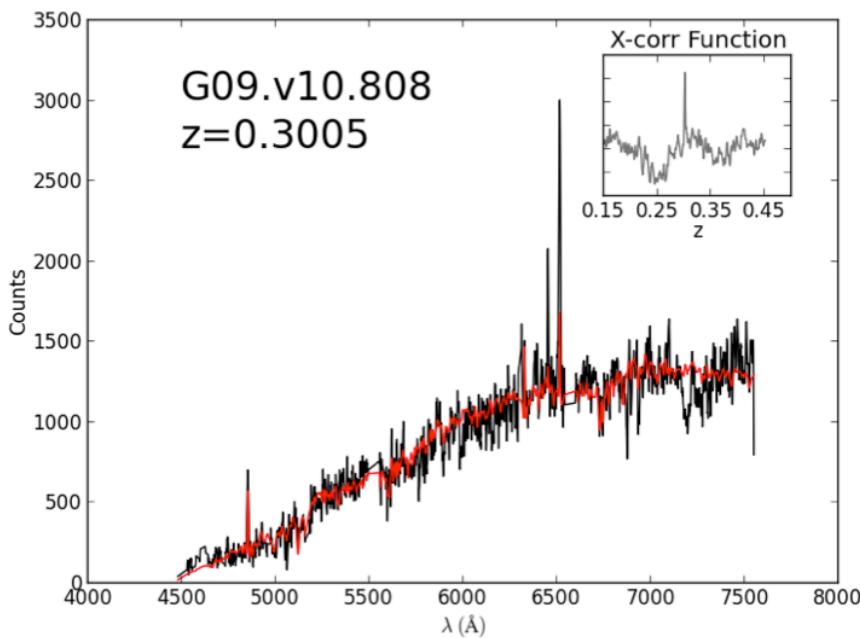
SALT Detected Spectra and Determined Redshifts



Leeuw, Crawford, H-ATLAS-team et al. SAIP 2013



SALT Detected Spectra and Determined Redshifts



Leeuw, Crawford, H-ATLAS-team et al. SAIP 2013

Implications of the SALT data

- The SALT spectra are confirming the foreground distance ($0.2 < z < 1$), first estimated from photometry, of the candidate optical lensing galaxies, constraining their redshifts by $z \pm 0.002$.
- Thus the data will constrain the lensing geometry, mass and magnification, with implications on constraining the physical properties of the lensed and lensing galaxies.
- About 1000 lenses are estimated to be discovered by the H-ATLAS survey, allowing samples large enough to help test galaxy evolutionary models and dark matter distribution and nature, by exploiting mass determinations of both baryonic and non-baryonic matter of the gravitational lenses.

Acknowledging Collaborators and Funders

- HATLAS-Core Team – Eales et al.
- HATLAS-Lensing Team – M. Negrello, S. Sergeant, S. Ambrose, A. Baker, L. Dunne, D. Riechers (**SALT Partner Members**)
- Steve Crawford, SALT and SAAO
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